

Isospin symmetry breaking in mirror nuclei

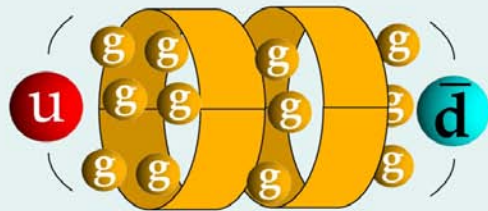
Silvia M. Lenzi

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Energy scales

Scales of energy in Nuclear Physics



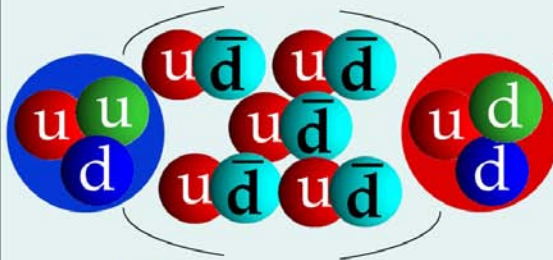
QCD scale



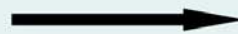
1000 MeV



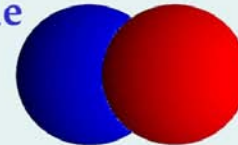
pion π^+
~140 MeV



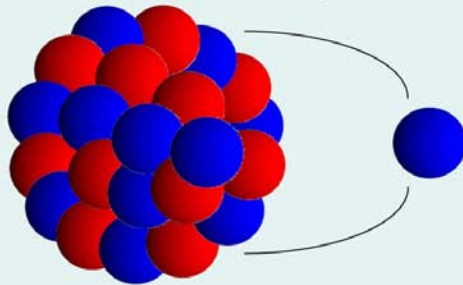
pion-mass scale



100 MeV



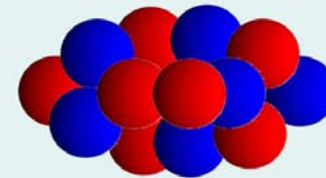
deuteron
~2 MeV



N-binding scale



10 MeV



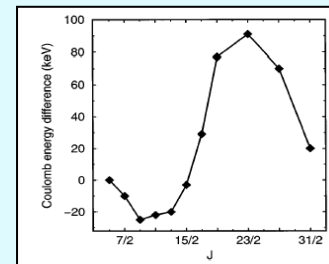
collective ~1 MeV



MED

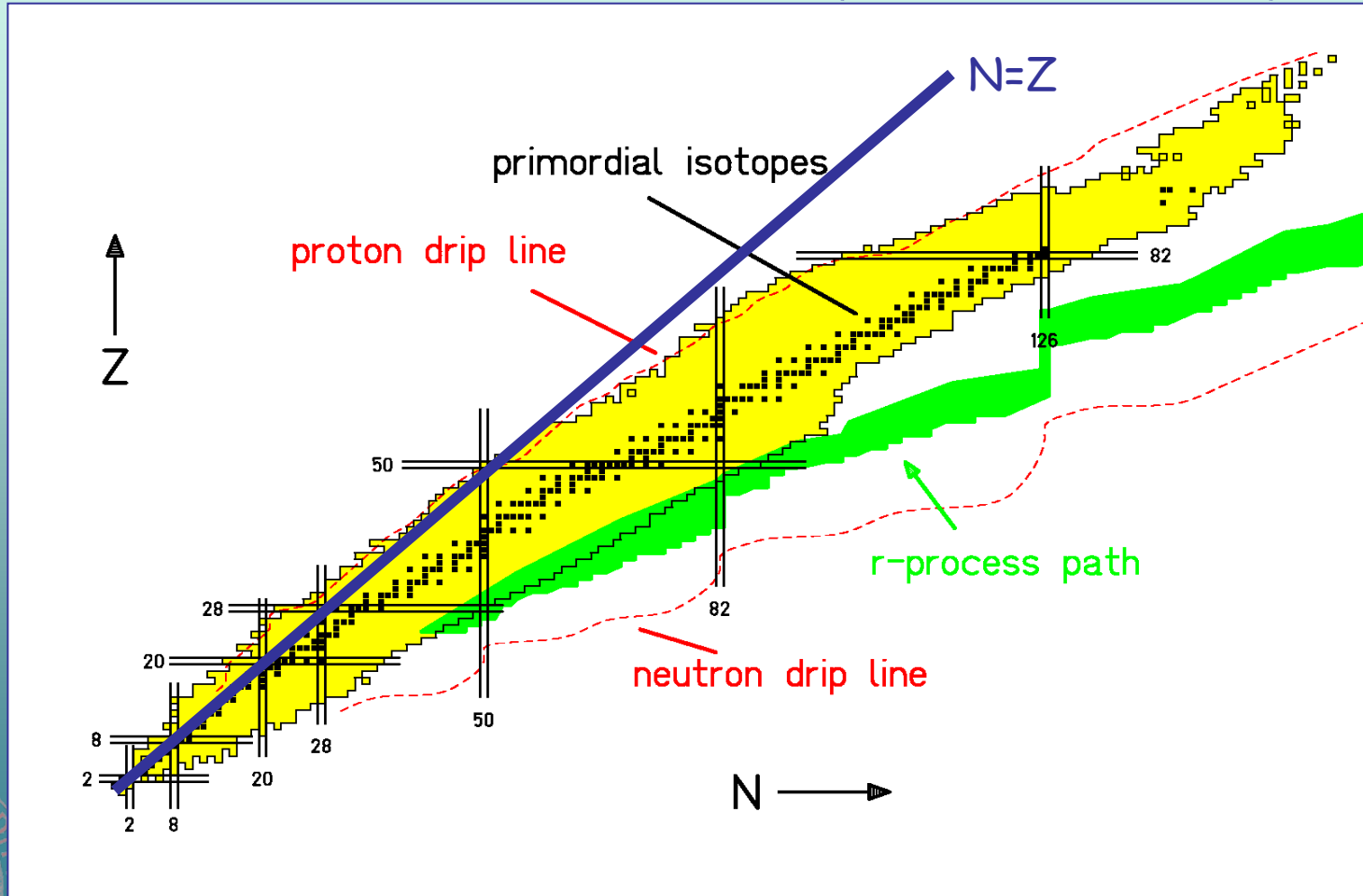


10-100 keV

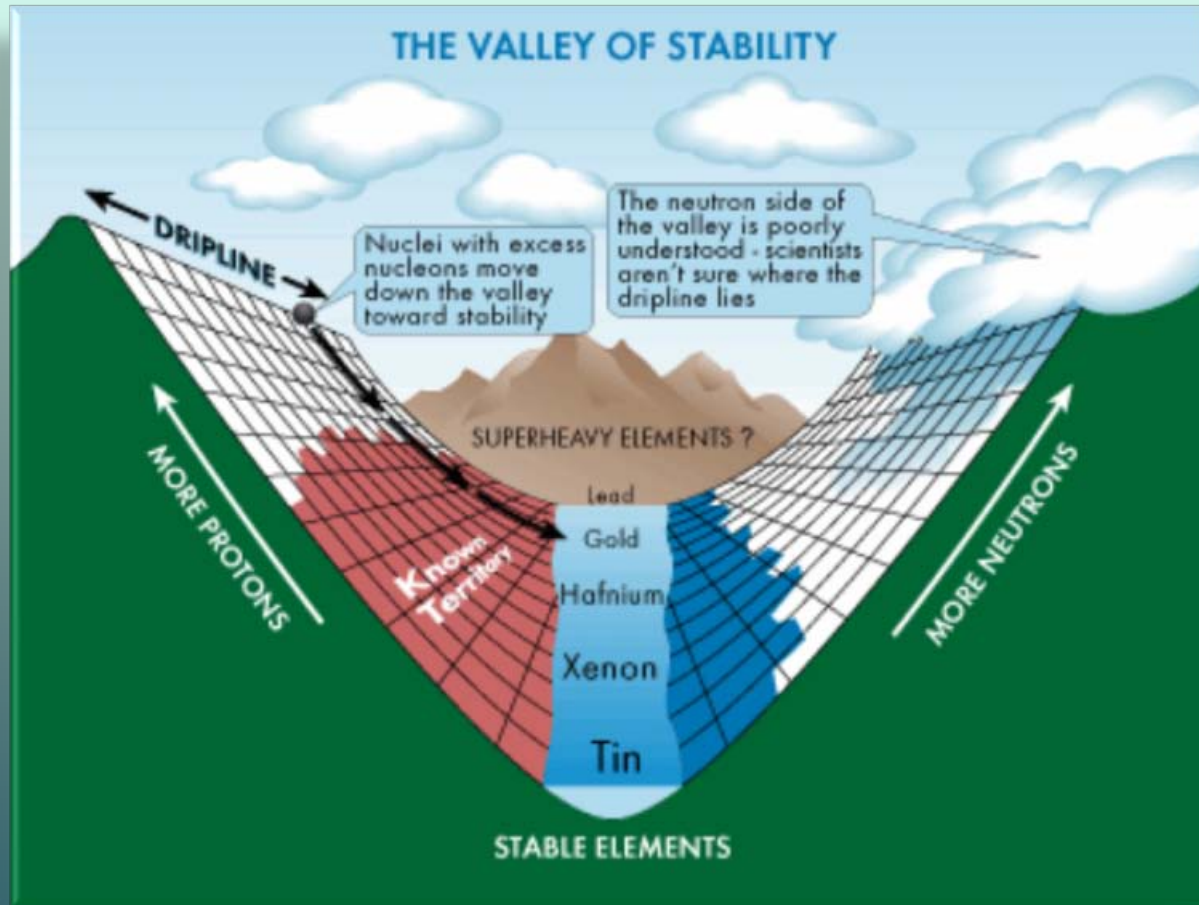


The Nuclear landscape

- 287 primordial isotopes exist in Nature
- ~6000 nuclei are bound (3000 are known)



A cut view

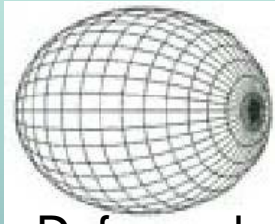
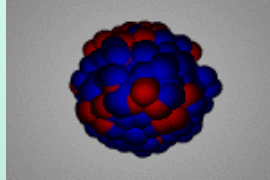
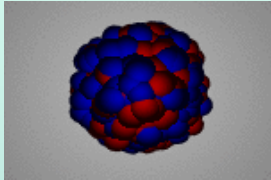


Istituto Nazionale
di Fisica Nucleare

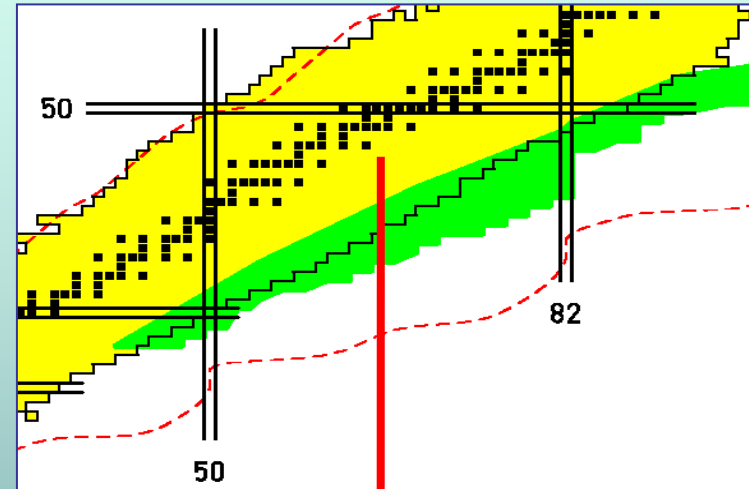
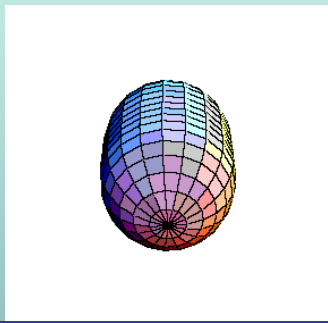
Nuclear shapes and excitation modes



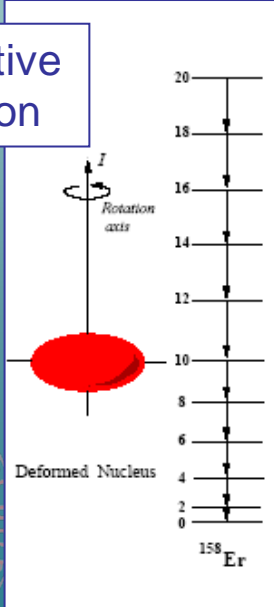
Spherical



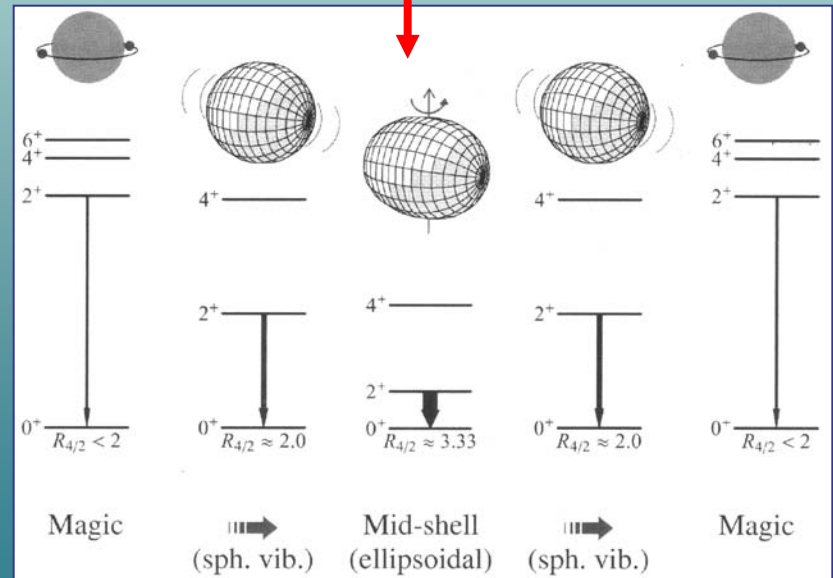
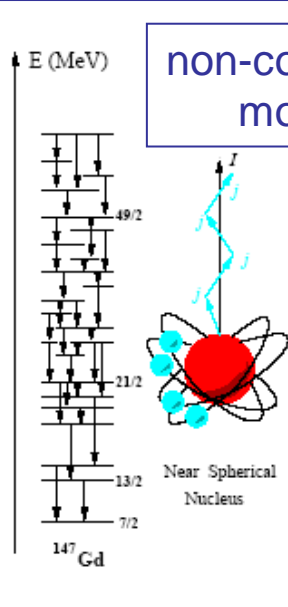
Deformed



collective motion

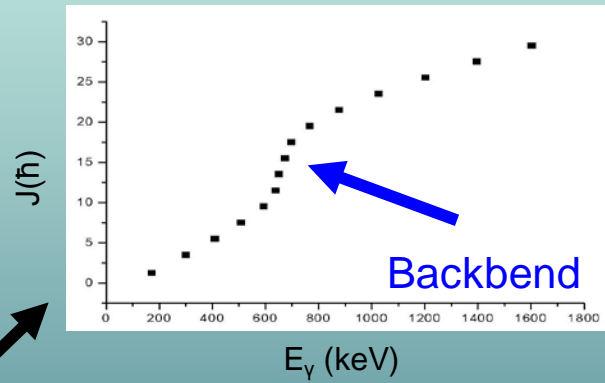
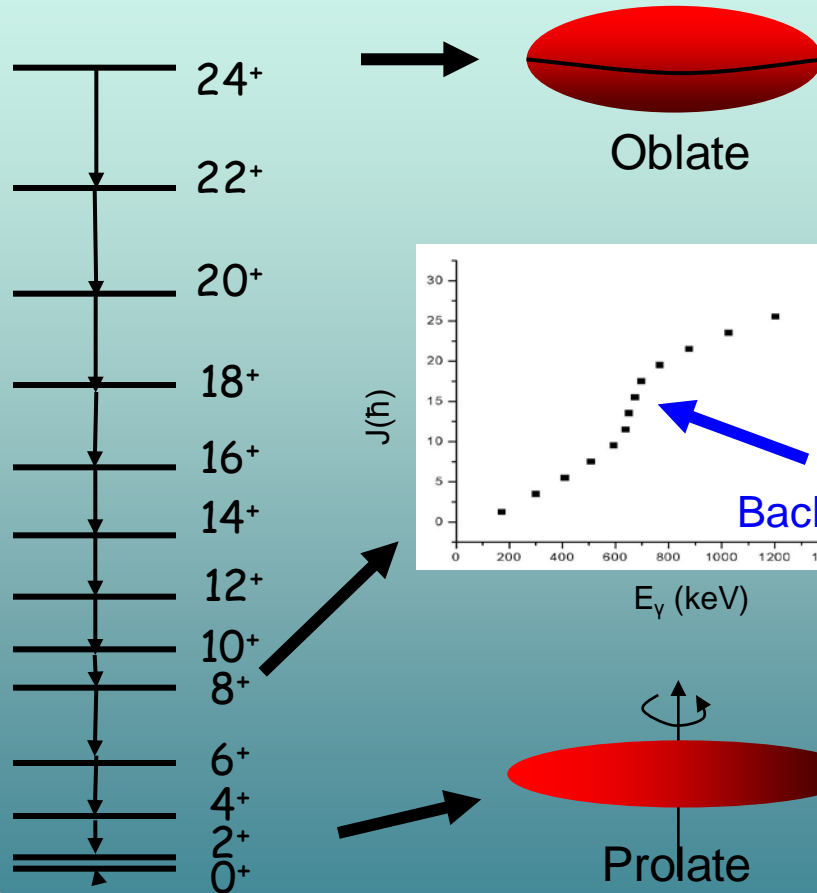


non-collective motion

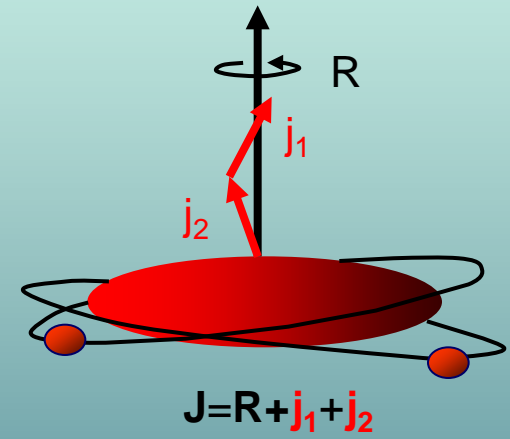


Changing deformation along the band

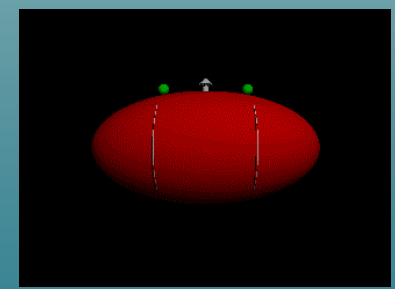
Rotational band



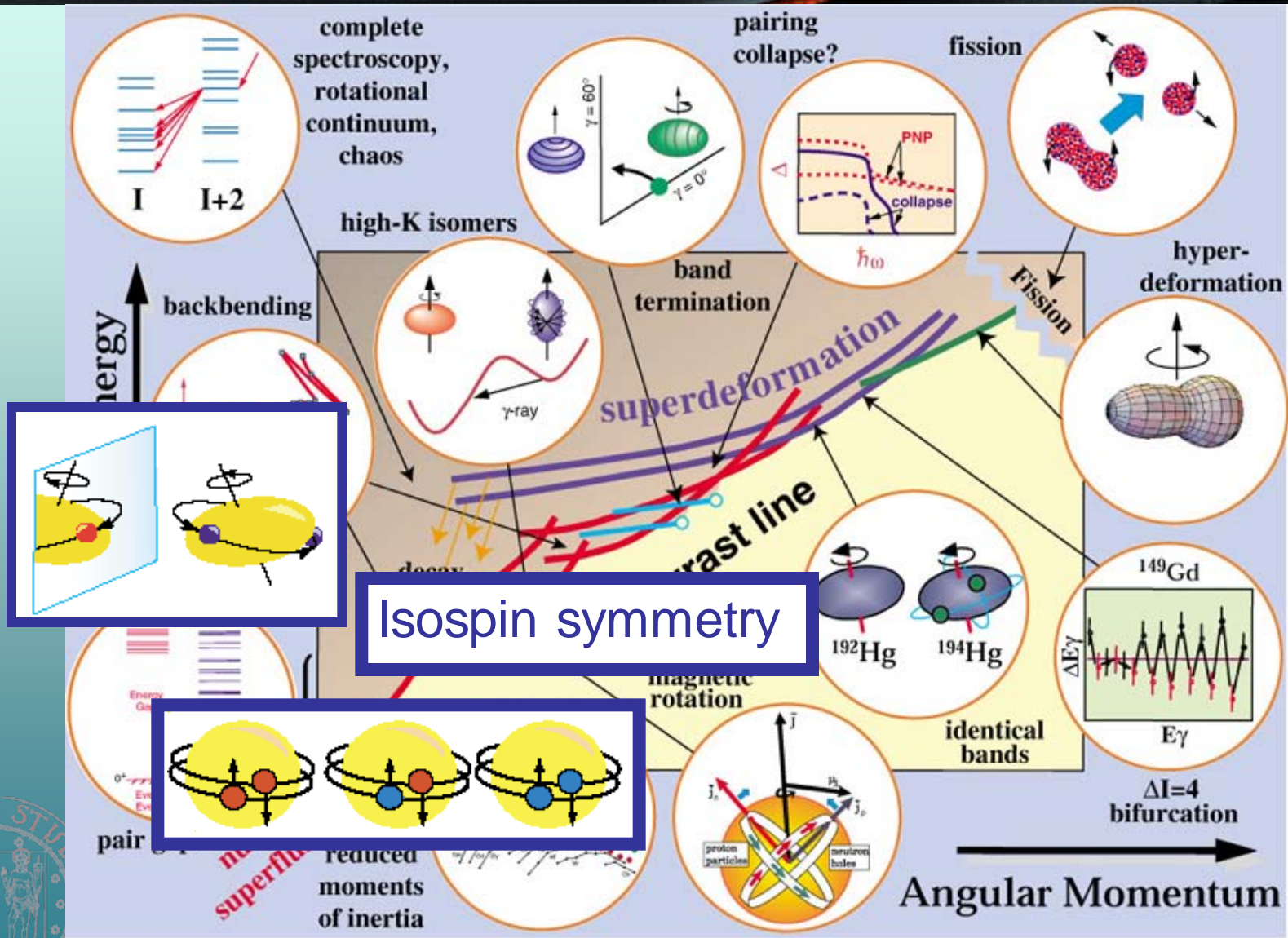
Band termination
all single spins are aligned



Deformed nucleus



The physics at the yrast line



Outline

What's isospin?

Why isospin symmetry?

How do we study isospin symmetry?

Experimental methods

Theoretical methods

What do we learn from the data?

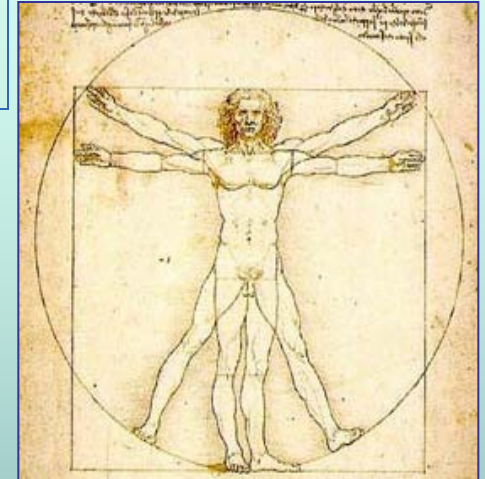
Background

What's isospin?
Why isospin symmetry?

Symmetries

Symmetries help to understand Nature

Examination of **fundamental symmetries**:
a key questions in Physics



↓

conservation laws

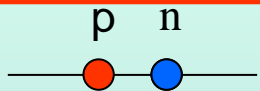
↓

good quantum numbers

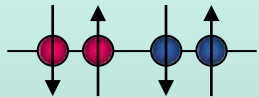
Conserved quantities imply some underlying symmetry enjoyed by the interactions

In Nuclear Physics symmetries help
to understand the behaviour of matter...

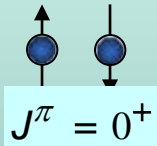
Symmetries in nuclear physics



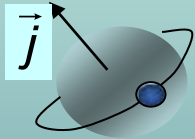
Isospin Symmetry: 1932 Heisenberg SU(2)



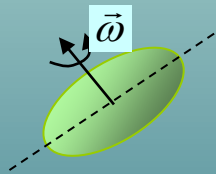
Spin-Isospin Symmetry: 1936 Wigner SU(4)



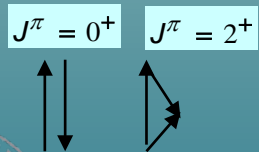
Seniority Pairing: 1943 Racah



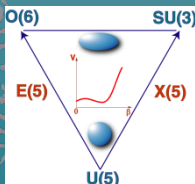
Spherical Symmetry: 1949 Mayer



Nuclear Deformed Field (spontaneous symmetry breaking)
Restore symm. → rotational spectra: 1952 Bohr-Mottelson
SU(3) Dynamical Symmetry: 1958 Elliott



Interacting Boson Model (IBM dynamical symmetry):
1974 Arima and Iachello

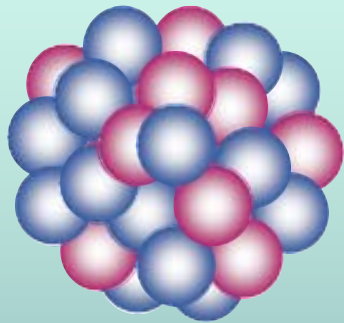


Critical point symm. E(5), X(5)
2000... F. Iachello

K. Heyde

The nucleus: a unique laboratory

Composed by two types of fermions differing only on its charge



Strong interaction: largely independent of the charge



Proton – Neutron exchange symmetry

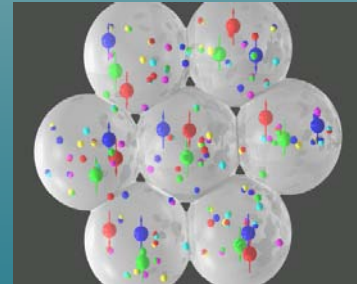
Proton and neutron can be viewed as two alternative states of the same particle: *the nucleon*.

The quantum number that characterizes the two charge states is the *isospin*

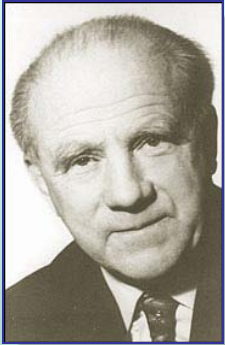
$$t = \frac{1}{2}$$

$$\begin{aligned} \text{proton} : \pi \\ t_z = -\frac{1}{2} \end{aligned}$$

$$\begin{aligned} \text{neutron} : \nu \\ t_z = +\frac{1}{2} \end{aligned}$$

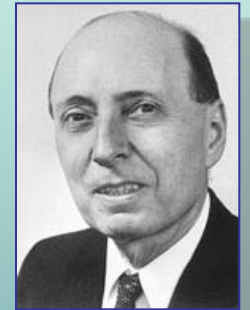


Charge symmetry and Isospin



1932 Heisenberg applies the Pauli matrices to the new problem of labeling the two alternative charge states of the nucleon.

1937 Wigner: isotopic spin is a good quantum number to characterize isobaric multiplets.

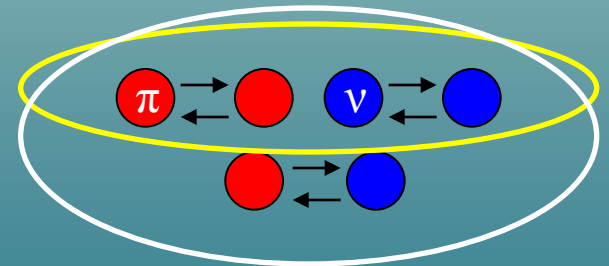


$$T_z = \sum_{i=1}^A t_{z,i} = \frac{N - Z}{2}$$

$$\left| \frac{N - Z}{2} \right| \leq T \leq \frac{N + Z}{2}$$



Isobaric analogue multiplets



- Charge Symmetry: $V_{pp} = V_{nn}$
- Charge Independence: $V_{pp} = V_{nn} = V_{pn}$



Two-nucleon system

For a two-nucleon system, four different isospin states can exist:

Triplet T=1

$$|T = 1, T_Z = 1\rangle = \uparrow\uparrow \quad |T = 1, T_Z = -1\rangle = \downarrow\downarrow \quad |T = 1, T_Z = 0\rangle = \frac{1}{\sqrt{2}} (\uparrow\downarrow + \downarrow\uparrow)$$

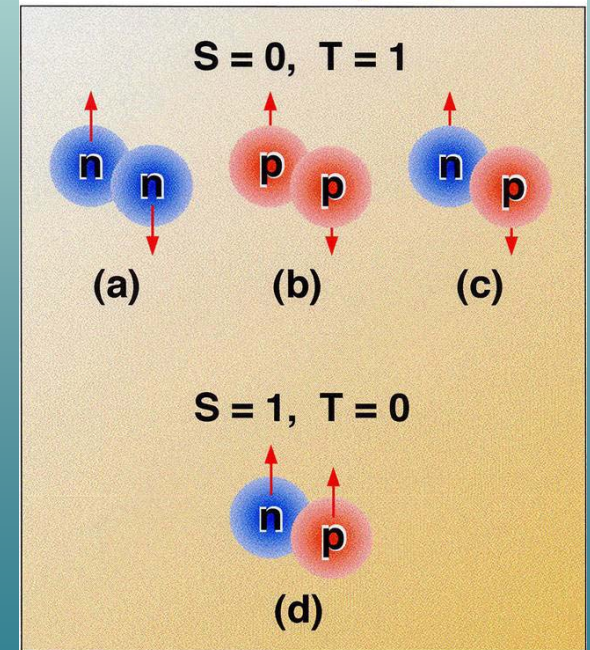
Singlet T=0

$$|T = 0, T_Z = 0\rangle = \frac{1}{\sqrt{2}} (\uparrow\downarrow - \downarrow\uparrow)$$

Total wavefunction must be antisymmetric on exchange of all (spin, space, isospin) co-ordinates...

$$\Psi_{\text{space}} \otimes \Psi_{\text{spin}} \otimes \Psi_{\text{isospin}}$$

nucleonic Cooper pairs



Isobaric multiplets

It is easy to extend the formalism to many-nucleon systems

The isospin quantum number T directly couples together the two effects of charge symmetry/independence and the Pauli principle

$$T_z = \sum_{i=1}^A t_{z,i} = \frac{N - Z}{2} \quad \left| \frac{N - Z}{2} \right| \leq T \leq \frac{N + Z}{2}$$

All nuclear states can be classified by the isospin quantum number, T , which to a good approximation can be treated as a good quantum number.

Since $T < T_z$ is forbidden, states of a given T can only occur in a set of nuclei with $T_z = T, T-1, \dots, -T \rightarrow$ isospin multiplet or “isobaric multiplet”.

This set of states of the same T in such a multiplet are termed “isobaric analogue states” (IAS).

Energy difference of ground states in A=27

Look at the difference in **absolute mass/binding energy**

Nuclei of same number
of particles A=27

<u>2.61 MeV</u>	<u>4.81 MeV</u>
^{27}Mg	^{27}Si
	<u>0</u>
	^{27}Al

Protons	12	13	14
Neutrons	15	14	13
<i>pp</i> -pairs	66	78	91
<i>nn</i> -pairs	105	91	78
<i>np</i> -pairs	180	182	182
<i>Total</i>	351	351	351

***pn*-pairs can exist in states NOT allowed for *pp* or *nn* pairs - PAULI PRINCIPLE**

^{27}Mg is less bound than ^{27}Al even though it has less protons,

but it has also less *p-n* T=0 pairs → T=0 stronger than T=1

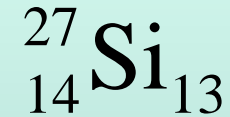
R.F. Casten



Analog states in isospin doublets

Mirror nuclei
with $T_z = \pm 1/2$

Test of
isospin
symmetry



	5.67	=====	9*				
5.44		=====		5*			5*
5.16	5.25	=====					3*
4.81							5*
4.51	4.58		7*				11*
	4.41						5*
3.96	4.05		1*				2*
	3.68						1*
2.98	3.00		9*				3*
	2.73						5*
2.21							7*
0.84	1.01						3*
							1*
							5*

stable

4.16s

T=1 isobaric triplets



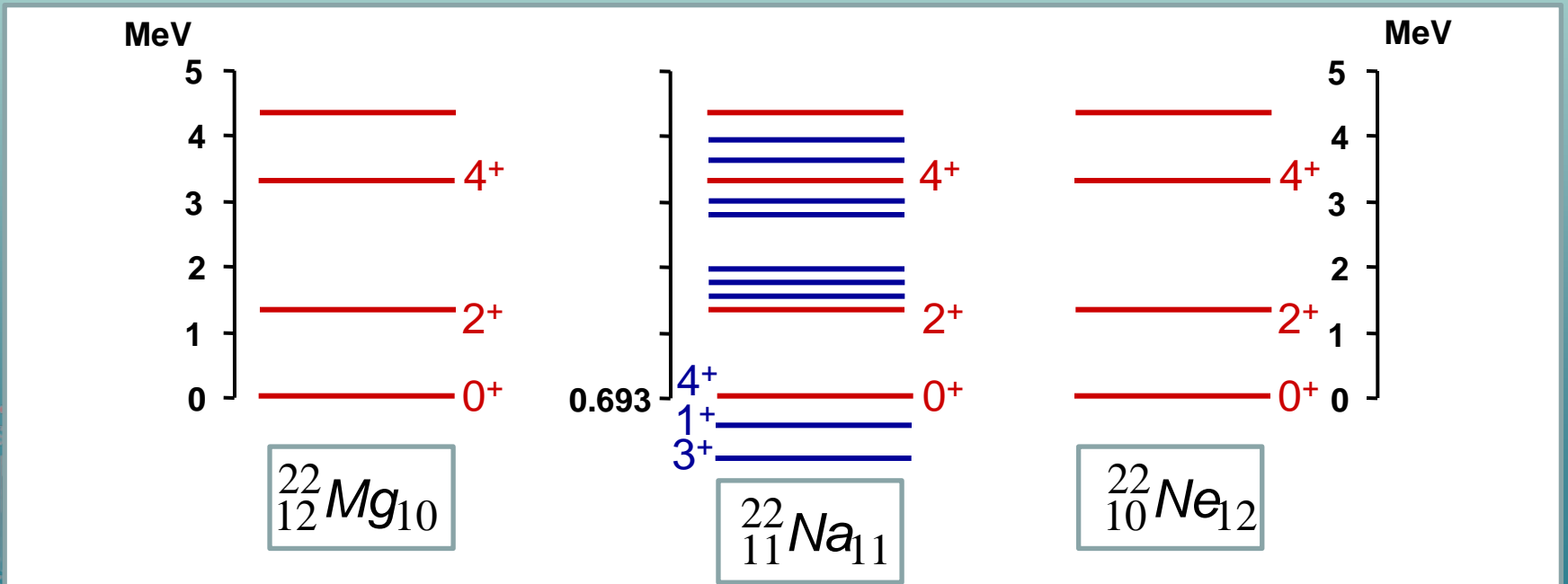
The nucleus can be characterized by isospin quantum numbers which restrict the possible states in which the many-nucleon system can exist.

Look at the isobaric triplet: ${}_{12}^{22}\text{Mg}_{10}$ ${}_{11}^{22}\text{Na}_{11}$ ${}_{10}^{22}\text{Ne}_{12}$

⇓ ⇑ ⇓ ⇑ ⇑

We expect: $T=1$ states low in energy in ${}^{22}\text{Mg}$ and ${}^{22}\text{Ne}$
 $T=0$ and $T=1$ states in ${}^{22}\text{Na}$ (N=Z)

Tests isospin independence

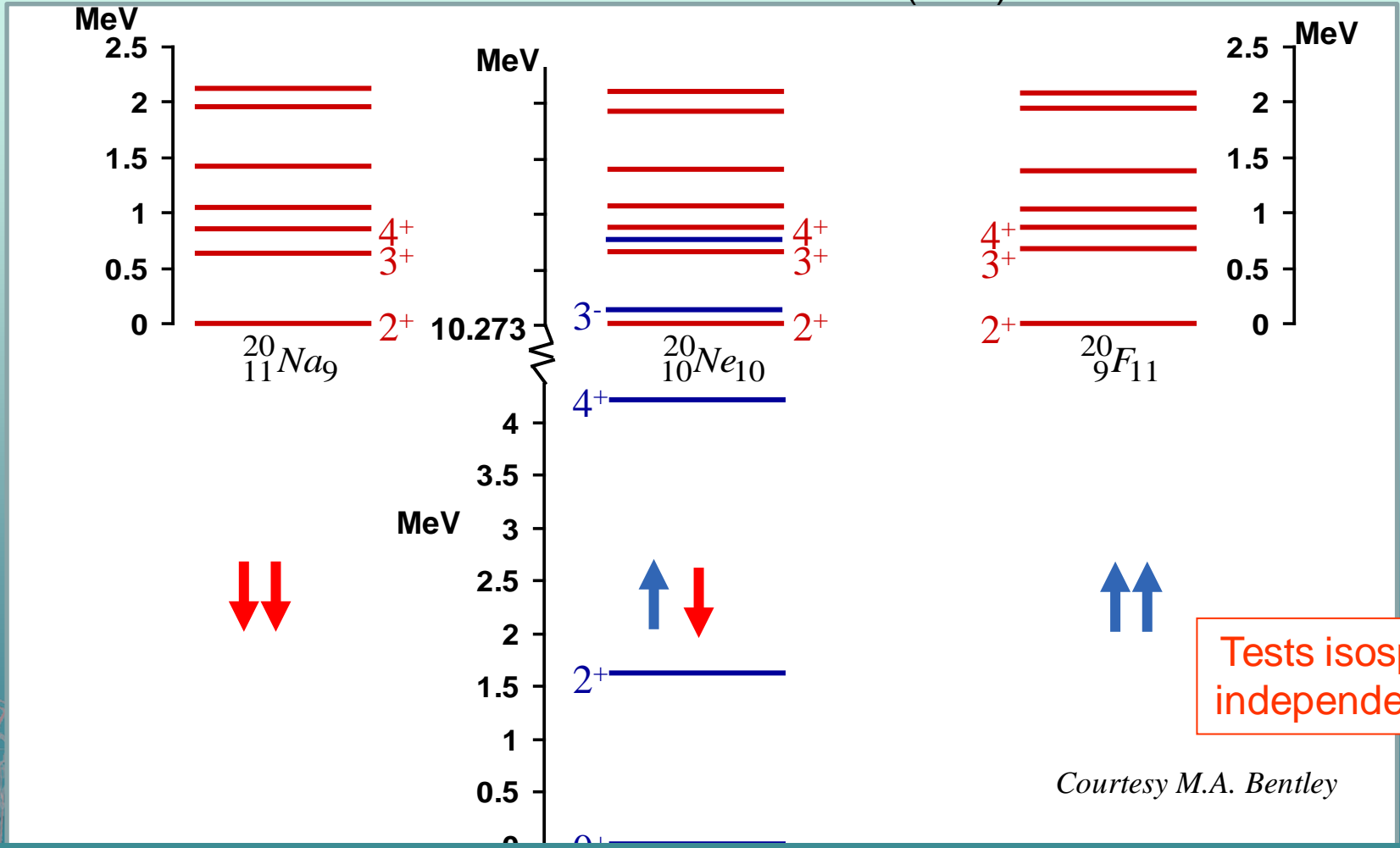


T=1 isobaric triplets



We expect: **T=1** states low in energy in ^{20}Na and ^{20}F

T=0 and **T=1** states in ^{20}Ne (N=Z)



Tests isospin independence

Courtesy M.A. Bentley

T ≥ 1 allowed

T ≥ 0 allowed

T ≥ 1 allowed



Three nucleon-system

We can form a doublet $T=1/2$

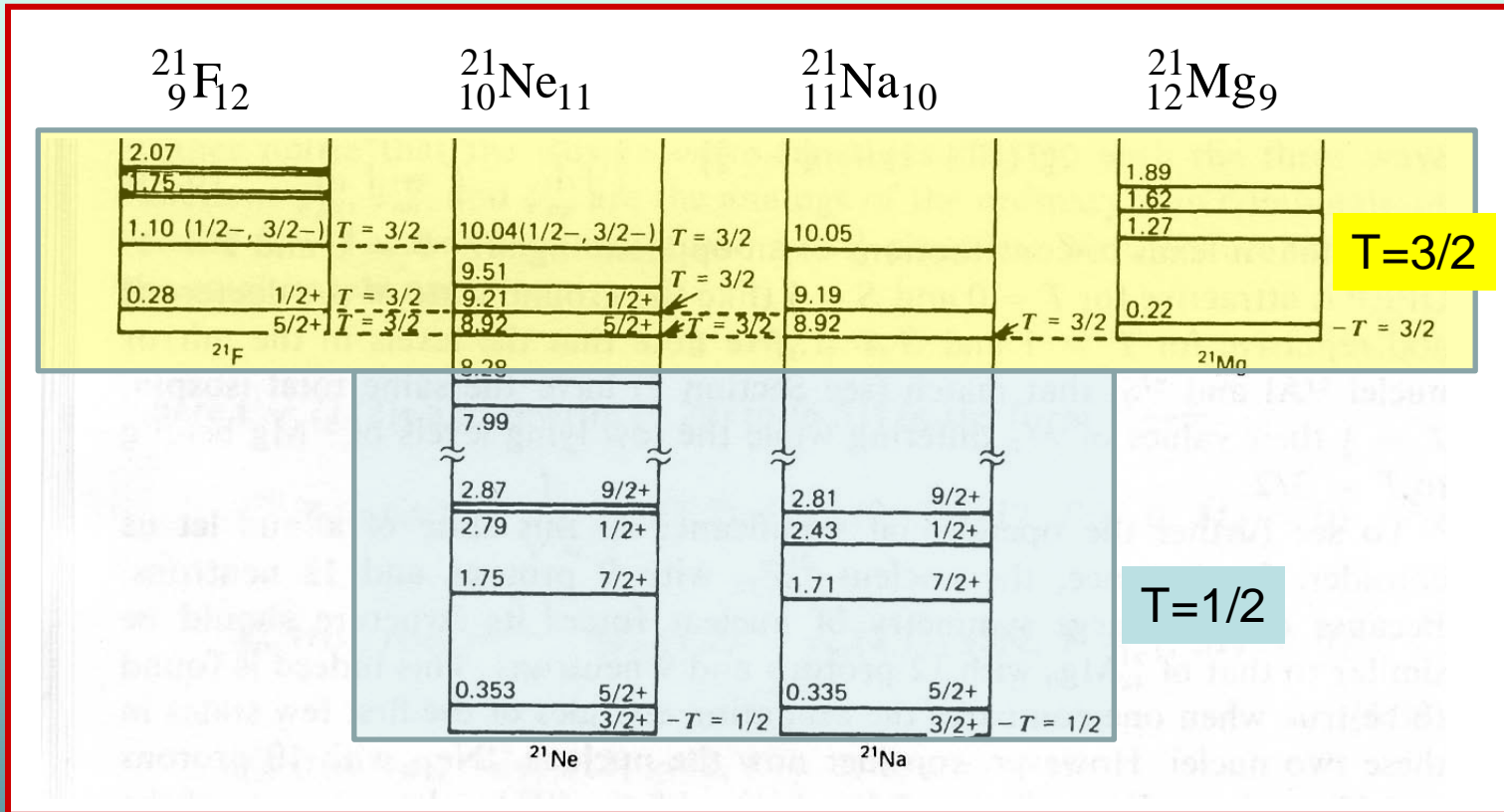
$$\left| T = \frac{1}{2}, T_z = +\frac{1}{2} \right\rangle = \uparrow\uparrow\downarrow \quad \left| T = \frac{1}{2}, T_z = -\frac{1}{2} \right\rangle = \uparrow\downarrow\downarrow$$

or a quadruplet $T=3/2$

$$\left| T = \frac{3}{2}, T_z = +\frac{3}{2} \right\rangle = \uparrow\uparrow\uparrow \quad \left| T = \frac{3}{2}, T_z = -\frac{3}{2} \right\rangle = \downarrow\downarrow\downarrow$$

$$\left| T = \frac{3}{2}, T_z = +\frac{1}{2} \right\rangle = \uparrow\uparrow\downarrow \quad \left| T = \frac{3}{2}, T_z = -\frac{1}{2} \right\rangle = \uparrow\downarrow\downarrow$$

T=3/2 Isobaric quadruplets: the spectra

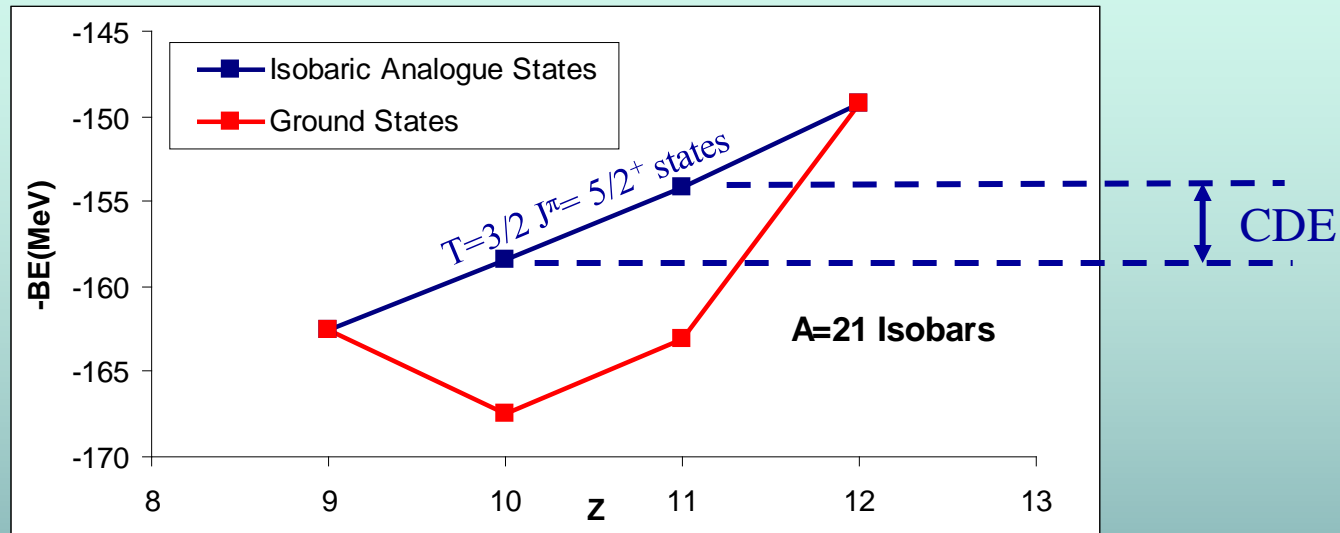


SMALL differences in excitation energy due mainly to Coulomb effects

LARGE differences in mass/binding energy - also due mainly to Coulomb effects....



Coulomb displacement energies (CDE) in A=21



CDE : For any two members of a multiplet of isospin T, transformed through exchange of k protons for neutrons is given by

$$(CDE)_{\alpha, T, T_z} = M_{\alpha, T, T_z + k} - M_{\alpha, T, T_z} + k\Delta M_{np}$$

(α = Isobaric analogue state, defined by A, T, J)



Isobaric Multiplet Mass Equation (IMME)

A simple relationship between mass and T_z for a set of analogue states (E.P.Wigner 1957)

To derive it we start with eigenstates of a charge-invariant (CI) Hamiltonian
The eigenvalues do not depend on T_z and analogue states are degenerate

$$H_{CI}|\alpha T T_z\rangle = E_{\alpha T}|\alpha T T_z\rangle$$

A charge-violating interaction lifts this degeneracy:

$$H_{CV} = \sum_{k=0}^2 H_{CV}^{(k)}$$

Any two-body force can be written in the isospin space in terms of 3 components

Isoscalar: $H_{CV}^{(0)} = (V_{nn} + V_{pp} + V_{pn})/3$

Isovector: $H_{CV}^{(1)} = V_{pp} - V_{nn}$

Isotensor: $H_{CV}^{(2)} = V_{pp} + V_{nn} - 2 V_{pn}$

IMME (2)

The total binding energy can be obtained as

$$BE(\alpha TT_z) = \langle \alpha TT_z | H_{CI} + H_{CV} | \alpha TT_z \rangle$$

The energy splitting in an isobaric multiplet can be obtained as:

$$\Delta BE(\alpha TT_z) = \langle \alpha TT_z | H_{CV} | \alpha TT_z \rangle$$

Let's apply the Wigner-Eckart theorem

$$\langle \alpha TT_z | H_{CV} | \alpha TT_z \rangle = \sum_{k=0}^2 (-1)^{T-T_z} \begin{pmatrix} T & k & T \\ -T_z & 0 & T_z \end{pmatrix} \langle \alpha T || H_{CV}^{(k)} || \alpha T \rangle$$

Calling

$$M^{(k)} = \langle \alpha T || H_{CV}^{(k)} || \alpha T \rangle$$

we obtain ...

IMME (3)

$$\begin{aligned} \Delta BE(\alpha T T_z) = & \frac{1}{\sqrt{2T+1}} M^{(0)} \\ & + \frac{T_z}{\sqrt{T(2T+1)(T+1)}} M^{(1)} \\ & + \frac{3T_z^2 - T(T+1)}{\sqrt{(2T-1)T(2T+1)(T+1)(2T+3)}} M^{(2)} \end{aligned}$$

We can now write the total BE into the form (a, b and c independent on T_z):

$$BE(\alpha T T_z) = a + b T_z + c T_z^2$$

Isoscalar, ~ 100's MeV
(includes the contribution of V_{CI})

Isovector ($V_{pp} \neq V_{nn}$)
~ 3-15 MeV (~ $A^{2/3}$)

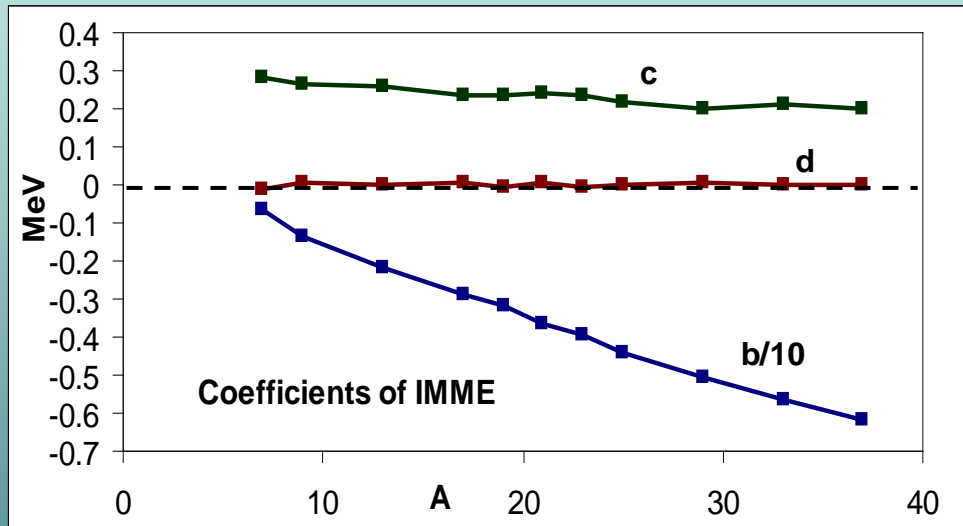
Isotensor ($V_{pp} + V_{nn} \neq 2V_{np}$)
~ 200-300 keV

The shift of the BE (mass) in an isobaric multiplet depends quadratically on T_z



Tests of IMME

IMME widely tested through fitting $T=3/2$ quadruplets. Coefficient of cubic term (dT_z^3) should be zero.



N. Auerbach Phys. Rep **98**(1983)273

Benenson & Kashy RMP **51**(1979)527

IMME works
beautifully!

The major contribution comes from the Coulomb interaction...

The Coulomb contribution to IMME

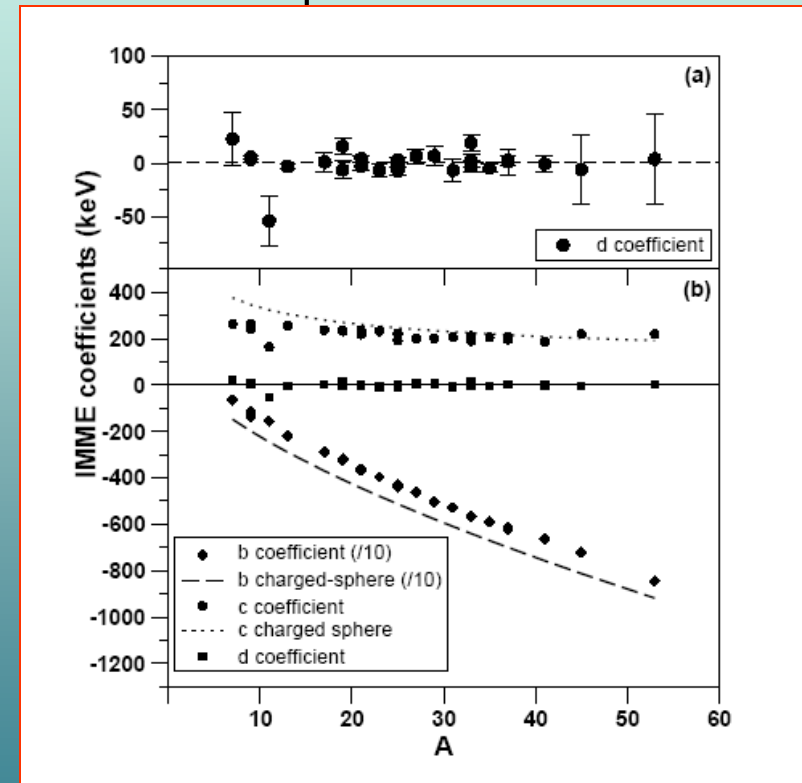
Consider the nucleus a uniform charged sphere, the Coulomb energy is:

$$E_C = \frac{3e^2 Z(Z-1)}{5R_C} = \frac{3e^2}{5r_0 A^{\frac{1}{3}}} \left[\frac{A}{4} (A-2) + (1-A)T_z + T_z^2 \right]$$

$$a = \frac{3e^2}{5r_0 A^{\frac{1}{3}}} \frac{A}{4} (A-2)$$

$$b = -\frac{3e^2}{5r_0 A^{\frac{1}{3}}} (A-1)$$

$$c = \frac{3e^2}{5r_0 A^{\frac{1}{3}}}$$



This crude estimate (dashed-line in the figure) has to be corrected by exchange terms (Pauli Principle), etc.

The Nolen-Schiffer Anomaly

IMME works well, but...reproducing magnitude of coefficients → historical problem

For two adjacent nuclei,

$$\begin{aligned} \text{CDE}_{\alpha, T, T_z} &= M_{\alpha, T, T_z} - M_{\alpha, T, T_z+1} + \Delta M_{np} \\ &= -b - c(2T_z + 1) + \Delta M_{np} \end{aligned}$$

Nolen & Schiffer (1969) calculated CDE for wide range of isobaric multiplets...

Used independent-particle models (with exchange term) → Corrected for other phenomena (e.g. **electromagnetic spin-orbit** (coming soon), magnetic effects, core-polarisation, isospin impurity, etc.).

But...magnitude of predicted CDE ALWAYS ~5-8% lower than experimental values. **Underestimated by ~500 keV!**

Auerbach (1983) improves the theoretical description but the anomaly remained

Duflo and Zuker (2002) reduce the difference to ~100-200 keV

Coulomb Energy Differences (CED)

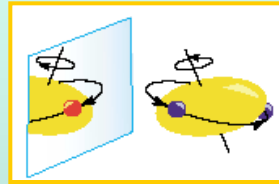
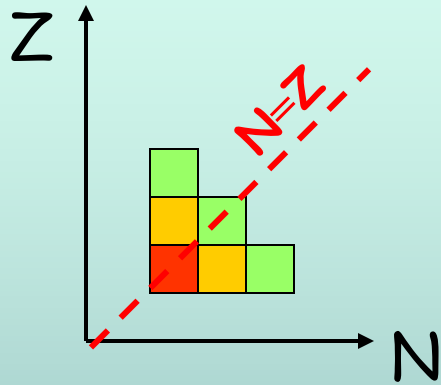
What about excited states - **Coulomb Energy Differences**...

“Normalise” ground state energies and take differences in excitation energy (CED)

How does Coulomb energy (+ other INC) change with E_x , J ?

Does Nolen-Schiffer Anomaly only concern ground states?

CED: MED and TED



Mirror Energy Differences

$$\text{MED}_J = E_{x_{J, T_z = -1/2}} - E_{x_{J, T_z = +1/2}} = -\Delta b_J$$

Test the charge symmetry of the interaction



Triplet Energy Differences

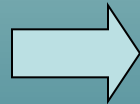
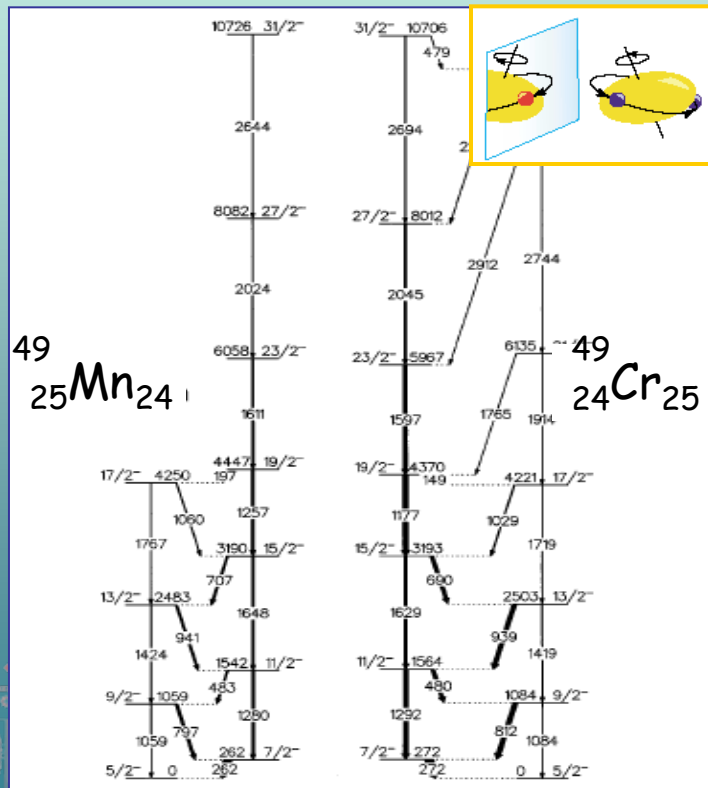
$$\text{TED}_J = E_{x_{J, T_z = -1}} + E_{x_{J, T_z = +1}} - 2E_{x_{J, T_z = 0}} = -2\Delta c_J$$

Test the charge independency of the interaction

A classical example: MED in T=1/2 states

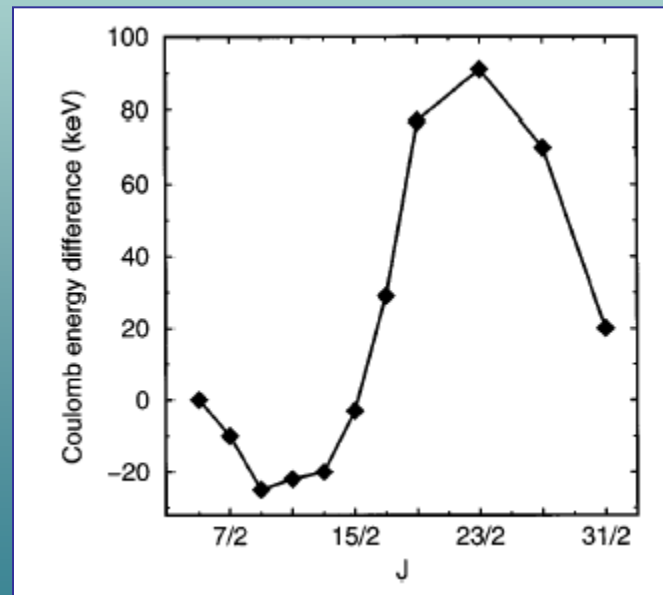
Coulomb effects in isobaric multiplets:

- bulk energy (100's of MeV)
- displacement energy (g.s.) CDE (10's of MeV)
- differences between excited states (10's of keV)

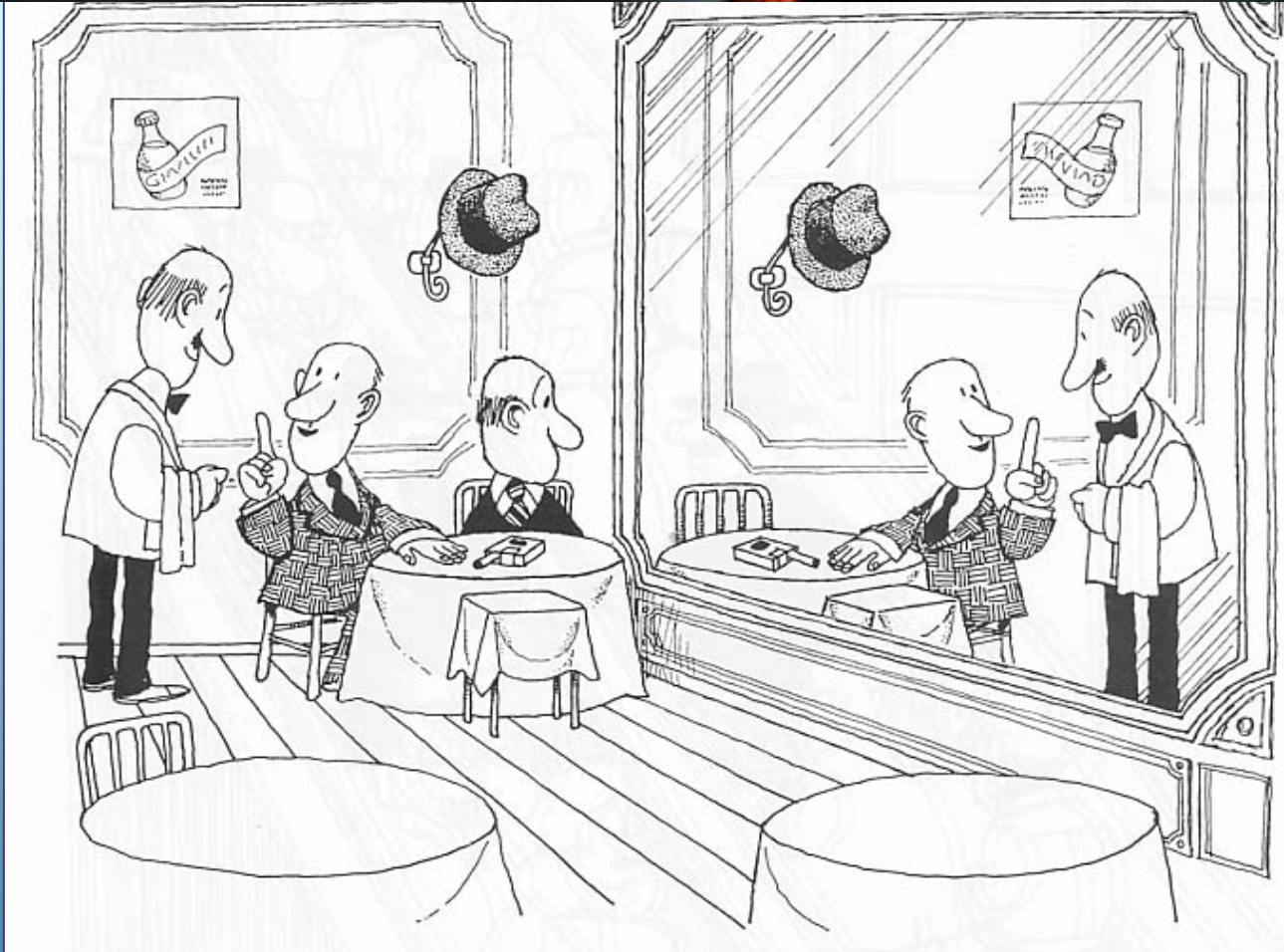


Mirror Energy Differences

$$MED_J = E_J(Z > N) - E_J(Z < N)$$



Symmetry breaking



Isospin symmetry breakdown, mainly due to the Coulomb field, manifests when comparing mirror nuclei. This constitutes an efficient observatory for a direct insight into nuclear structure properties.

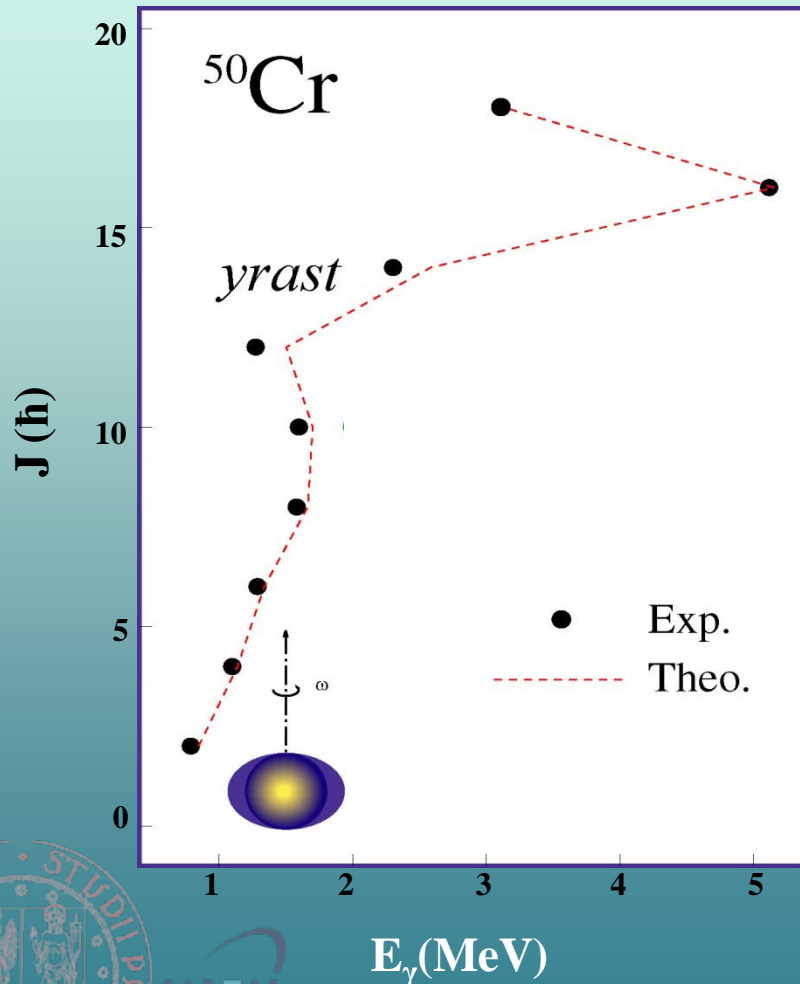
Measuring the Isospin Symmetry Breaking

We will see that CED are extremely sensitive to quite subtle nuclear structure phenomena which, with the aid of shell model calculations, can be interpreted quantitatively at the level of 10's of keV.

We measure *nuclear* structure features:

- How the nucleus generates its angular momentum
- Evolution of the radii (deformation) along a rotational band
- Learn about the configuration of the states
- Isospin non-conserving terms in the nuclear interaction

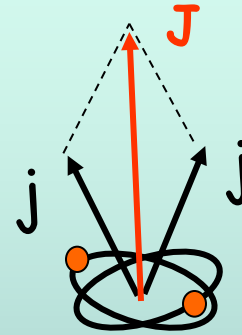
Understanding structure features



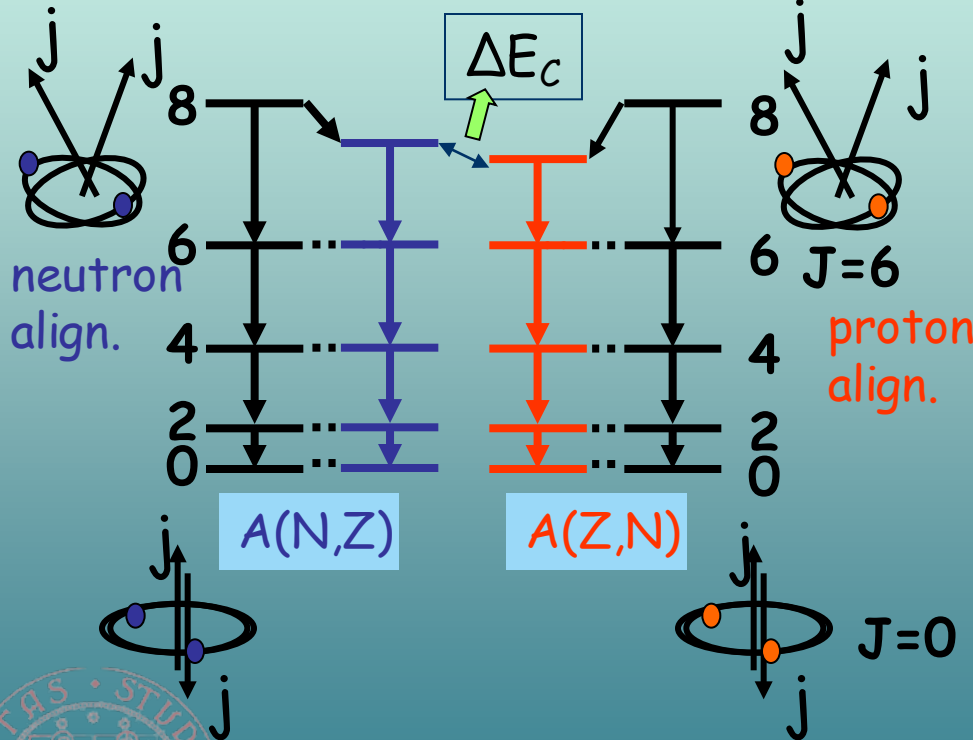
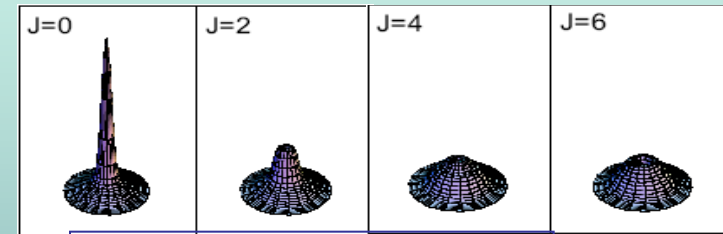
Most of the structure features of nuclei in the $f_{7/2}$ shell are very well described by shell model calculations in the full fp valence space

- What happens at the backbending?
- band-crossing?
 - alignment?
 - which nucleons are aligning?

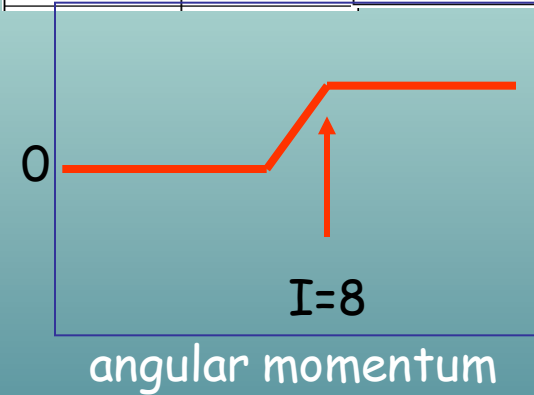
Mirror energy differences and alignment



probability distribution for the relative distance of two like particles in the $f_{7/2}$ shell



MED



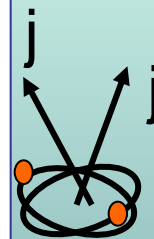
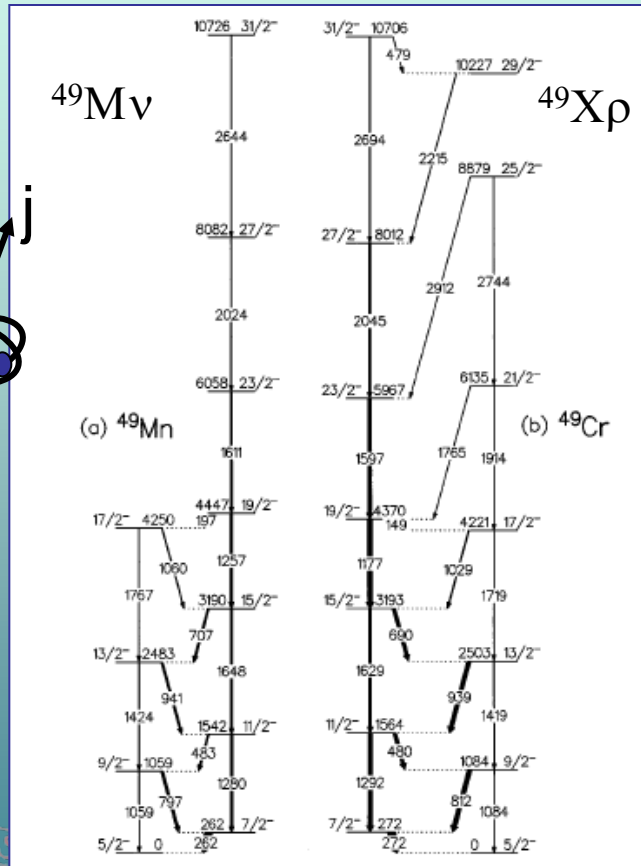
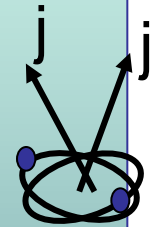
Shifts between the excitation energies of the mirror pair at the back-bend indicate the type of nucleons that are aligning

J.A. Cameron *et al.*,
Phys. Lett. B **235**, 239 (1990)



Nucleon alignment at the backbending

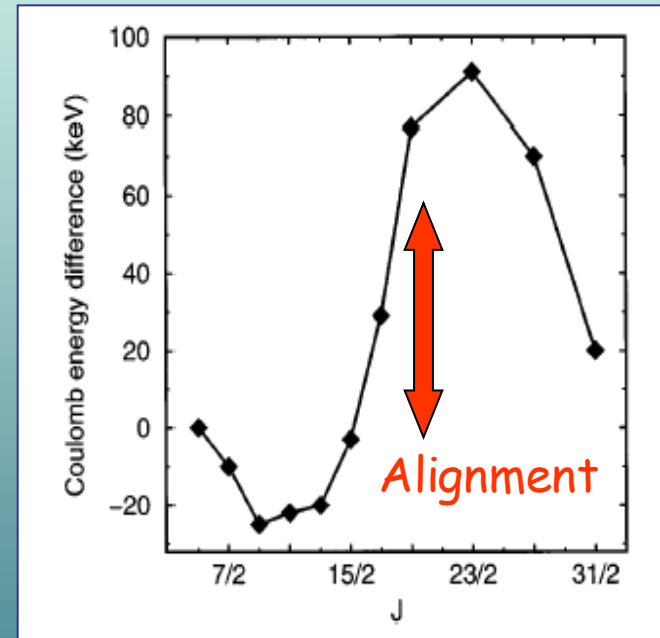
J.A. Cameron *et al.*, Phys. Lett. **B 235**, 239 (1990)
 C.D. O'Leary *et al.*, Phys. Rev. Lett. **79**, 4349 (1997)



$J=6$



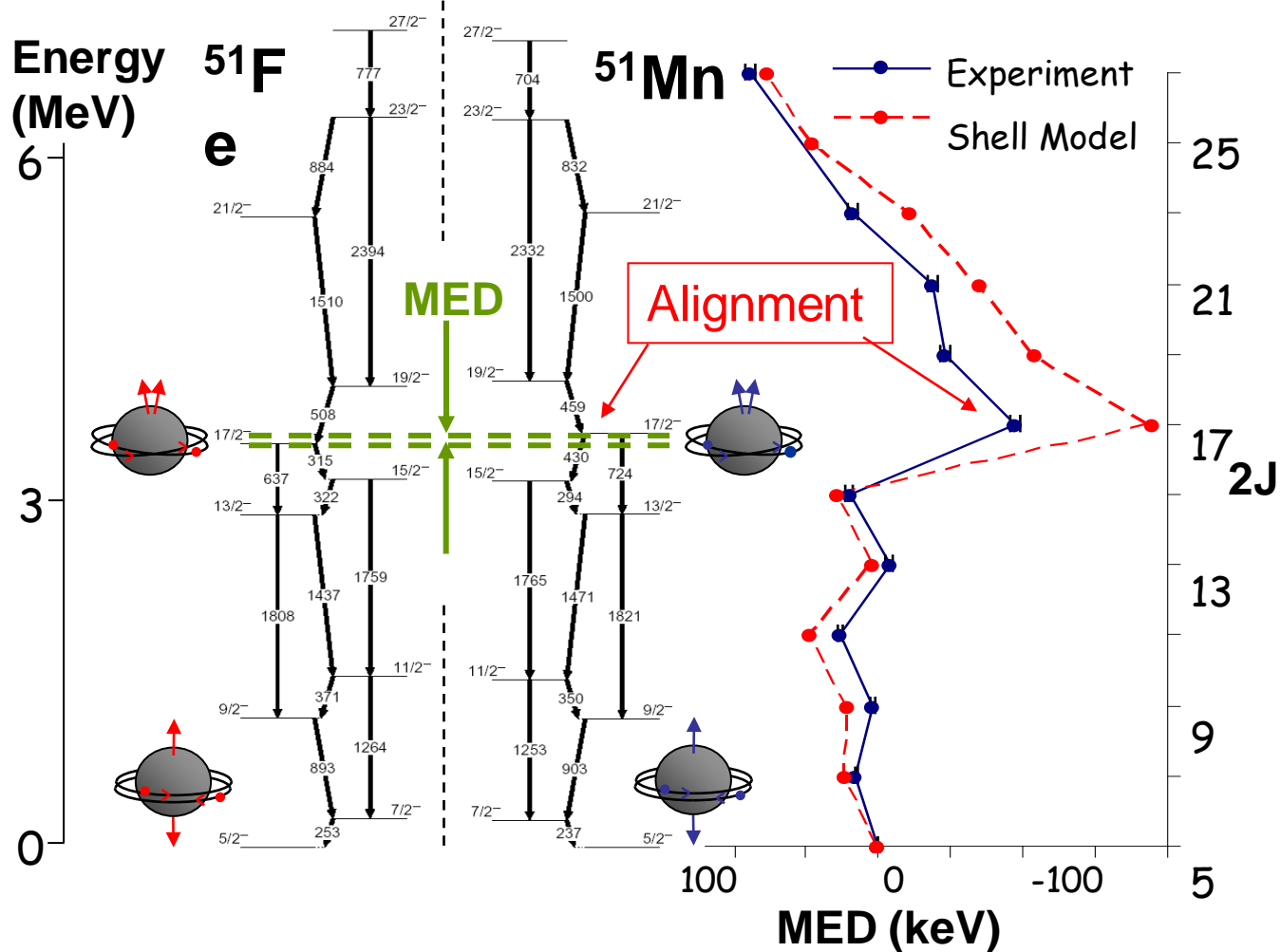
Experimental MED




MED are a probe of nuclear structure:
 reflect the way the nucleus generates its angular momentum

Nucleon alignment at the backbending

D.D. Warner, M.A. Bentley and P. Van Isacker., Nature Physics 2 (2006) 311



Cranked shell model and alignment



ELSEVIER

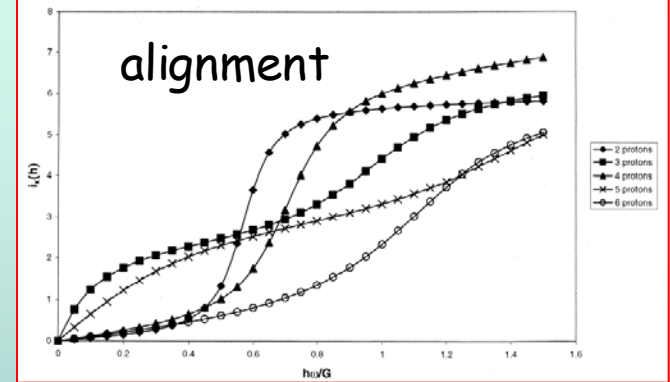
10 December 1998

PHYSICS LETTERS B

Physics Letters B 443 (1998) 16–20

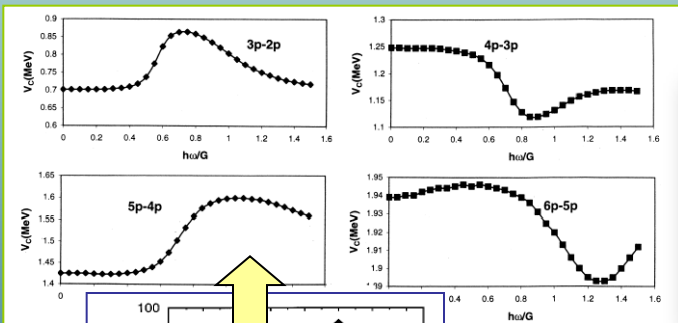
Rotational dependence of Coulomb energy differences

J.A. Sheikh ^a, D.D. Warner ^b, P. Van Isacker ^c



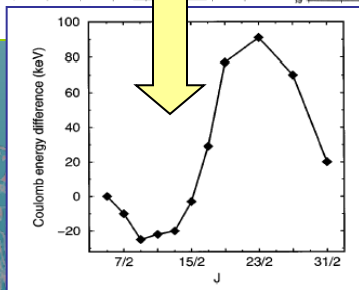
Cranked shell-model

$$H' = h_{def} + V_2 - \omega J_x$$

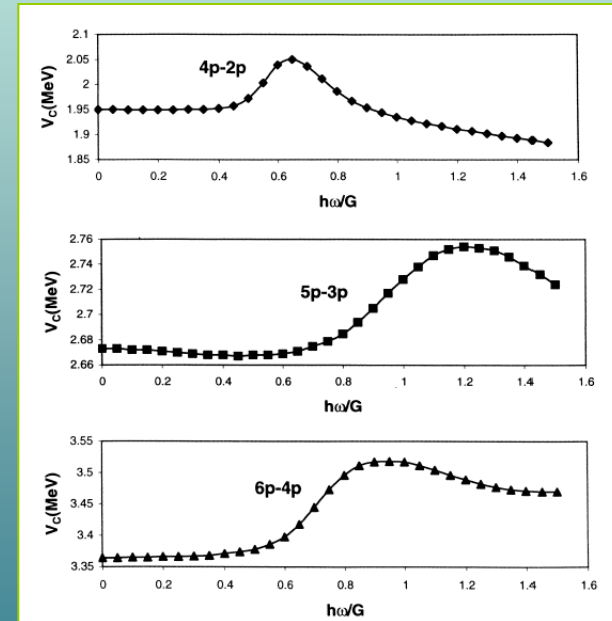


Approximations:

- one shell only
- fixed deformation
- no p-n pairing



CSM: good qualitative description of the data



J.A. Sheikh, P. Van Isacker, D.D. Warner and J.A. Cameron,
Phys. Lett. B 252 (1990) 314



Alignment and shell model

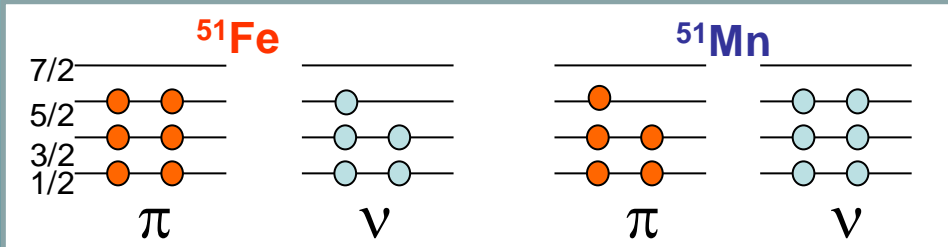
Define the operator

$$\mathbf{A}_\pi = \left[\left(a_\pi^+ a_\pi^+ \right)^{J=6} \left(a_\pi a_\pi \right)^{J=6} \right]^0$$

“Counts” the number of protons coupled to J=6

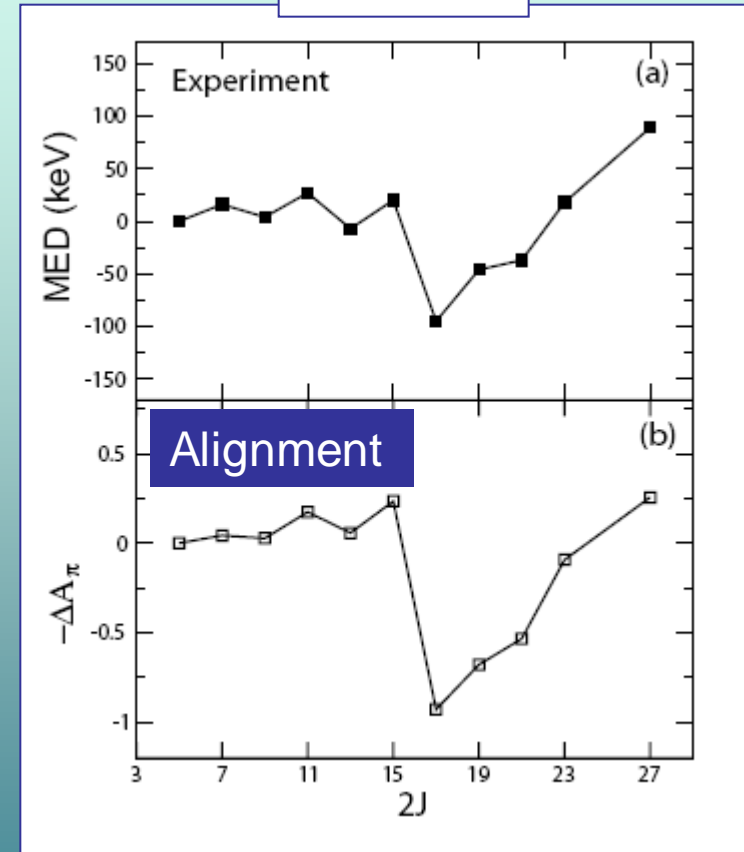
Calculate the difference of the expectation value in both mirror as a function of the angular momentum

$$\Delta \mathbf{A}_{\pi, J} = \langle \Phi_J | \mathbf{A}_\pi(Z_>) | \Phi_J \rangle - \langle \Phi'_J | \mathbf{A}_\pi(Z_<) | \Phi'_J \rangle$$



In ^{51}Fe (^{51}Mn) a pair of **protons** (**neutrons**) align first and at higher frequency align the **neutrons** (**protons**)

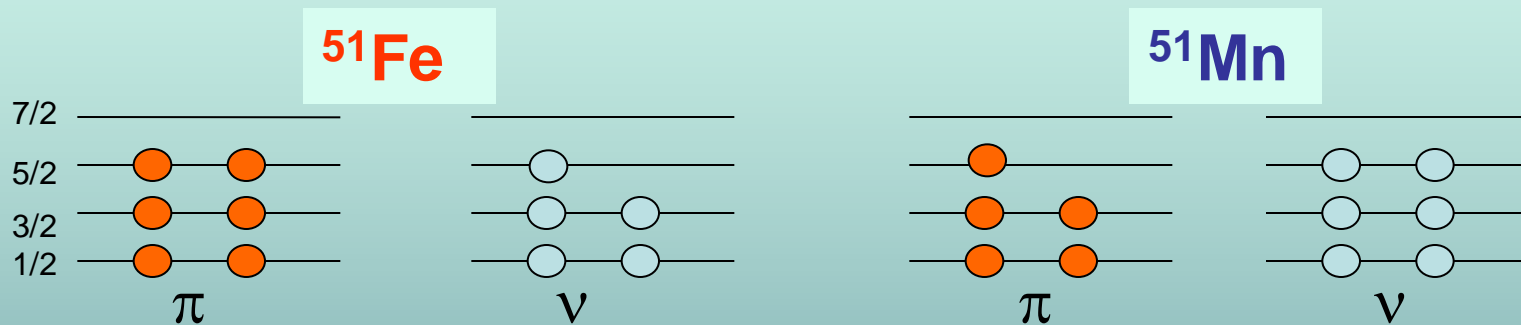
^{51}Fe - ^{51}Mn



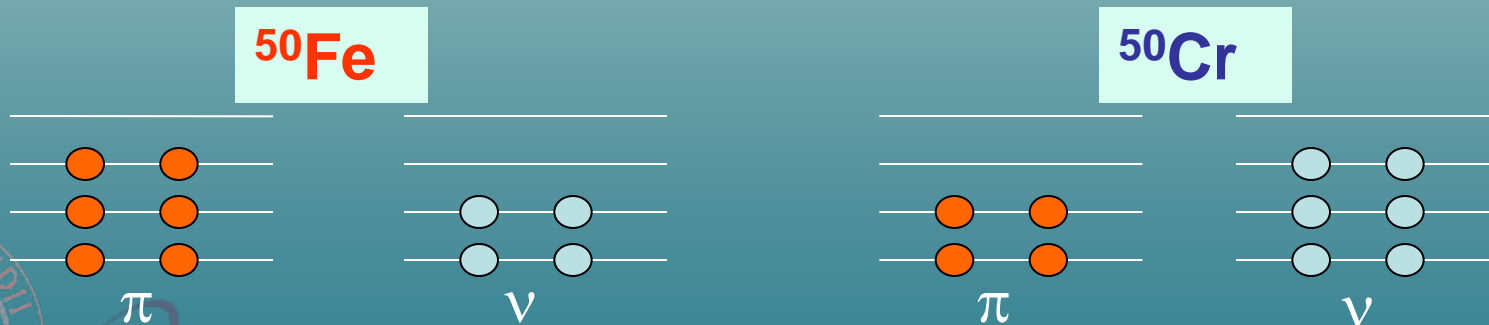
M.A.Bentley et al. Phys Rev. C62 (2000) 051303

Alignment in odd- and even-mass nuclei

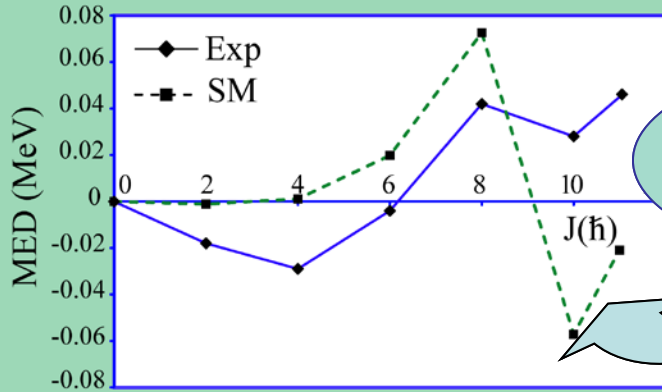
In **odd-mass nuclei**, the type of nucleons that aligns first is determined by the **blocking effect** → the even fluid will align first



What about **even-even** rotating nuclei?



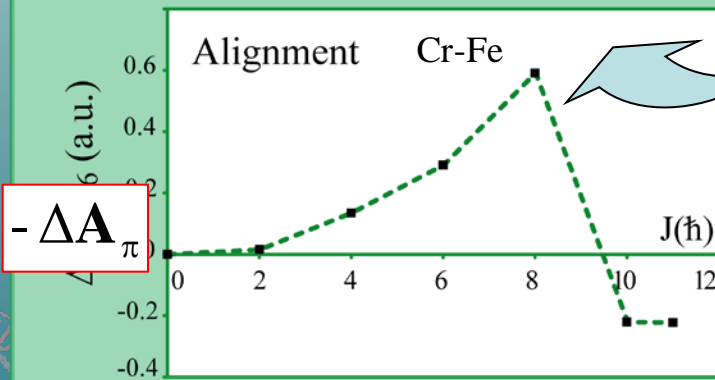
Alignment in even-even rotating nuclei



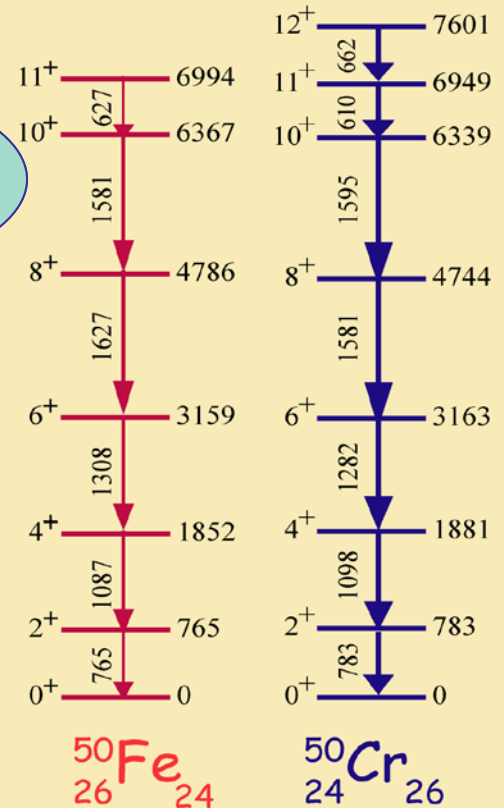
Renormalization of the Coulomb m.e.? Only a Coulomb effect?

calculate for protons in both mirrors:

$$\Delta A_\pi = \langle \Phi_J | A_\pi(Z_>) | \Phi_J \rangle - \langle \Phi'_J | A_\pi(Z_<) | \Phi'_J \rangle$$



counts the number of aligned protons



In ^{50}Cr (^{50}Fe) a pair of protons (neutrons) align first and at higher frequency align the neutrons (protons)

can shell model do better?

S.M.Lenzi et al., Phys. Rev. Lett. 87, 122501 (2001)



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- 3) Review article on CDE: J.A. Nolen and J.P. Schiffer, Ann. Rev. Nucl. Sci 19 (1969) 471
- 4) N. Auerbach, Phys. Rep. 98 (1983) 273
- 5) Theory on CDE: J. Duflo and A.P. Zuker, Phys. Rev. C 66 (2002) 051304(R)
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- 7) Theory on CED: J.A. Sheikh, D.D. Warner and P. Van Isacker, Phys. Lett. B 443 (1998) 16
- 8) Shell model reviews: B.A. Brown, Prog. Part. Nucl. Phys 47 (2001) 517; T. Otsuka, M. Honma, T. Mizusaki and N. Shimitzu, Prog. Part. Nucl. Phys 47 (2001) 319; E. Caurier, G. Martinez-Pinedo, F. Nowacki, A.Poves, and A.P. Zuker, Rev. Mod. Phys. 77 (2005) 427
- 9) Review article on $N \sim Z$: D. D. Warner, M. A. Bentley, P. Van Isacker, Nature Physics 2, (2006) 311 - 318



Lecture 1

The end

